

# libRadtran user course, lecture # 7

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# Radiation quantities

Solution of the radiative transfer equation generally yields the **diffuse radiance**

$$L(\tau, \mu, \phi)$$

and the **direct radiance**

$$L^{\text{dir}}(\tau, \mu_0, \phi_0).$$

For polarization above quantities are vectors. From the radiances the upward,  $E^\uparrow(\tau)$ , and downward,  $E^\downarrow(\tau)$ , fluxes, or **irradiances** [ $\text{W m}^{-2} \text{ nm}^{-1}$ ], are calculated

$$E^\uparrow(\tau) = \int_0^{2\pi} d\phi \int_0^1 \mu L(\tau, \mu, \phi) d\mu$$

$$E^\downarrow(\tau) = \mu_0 L_0 e^{-\tau/\mu_0} + \int_0^{2\pi} d\phi \int_0^1 \mu L(\tau, -\mu, \phi) d\mu.$$

# Radiation quantities cont'd

Furthermore, the mean intensity

$$\overline{L(\tau)} = \frac{1}{2\pi} \left[ L_0 e^{-\tau/\mu_0} + \int_0^{2\pi} d\phi \int_0^1 L(\tau, -\mu, \phi) d\mu + \int_0^{2\pi} d\phi \int_0^1 L(\tau, \mu, \phi) d\mu \right],$$

is related to the actinic flux (Madronich, 1987),  $F$ , used for the calculation of photolysis (or photodissociation) rates

$$F(\tau) = 4\pi \overline{L(\tau)}.$$

Finally, heating rates [K/day] may be calculated from either the flux differences or the mean intensity.

$$\frac{\partial T}{\partial t} = \frac{1}{c_p \rho_m} \frac{\partial E_{\text{net}}}{\partial z} = \frac{4\pi}{c_p \rho_m} (1 - \omega)(\overline{L} - B) \frac{\partial \tau}{\partial z}.$$

Note that the partial derivative of  $\tau$  with respect to  $z$  is needed since optical properties and  $\overline{L}$  are calculated as functions of  $\tau$ .

# Radiation quantities cont'd

**Albedo:**

$$A(\lambda) = \frac{E^\uparrow(\lambda)}{E^\downarrow(\lambda)}$$

**Global albedo:**

$$A_g = \frac{\int_0^\infty E^\uparrow(\lambda) d\lambda}{\int_0^\infty E^\downarrow(\lambda) d\lambda},$$

**Radiative forcing (IPCC):**

*Radiative forcing is a measure of the influence a factor has in altering the balance of incoming and outgoing energy in the Earth-atmosphere system and is an index of the importance of the factor as a potential climate change mechanism.*

Includes the calculation of the change in net (down minus up) irradiance (solar plus longwave [ $\text{W m}^{-2}$ ]) due to the factor.

**Brightness temperature**  $T$  (Planck's law, black body radiation)

$$L(\lambda, T) = \frac{2hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda kT}} - 1}$$

# Measured quantities, NILU-UV

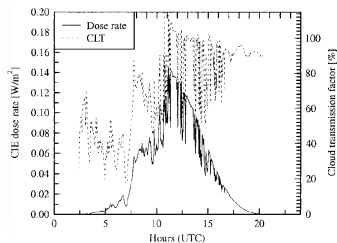
Measures irradiances in band  $i$

$$E_i^\downarrow(t) = \int_{\lambda_1}^{\lambda_2} E^\downarrow(\lambda, t) d\lambda$$

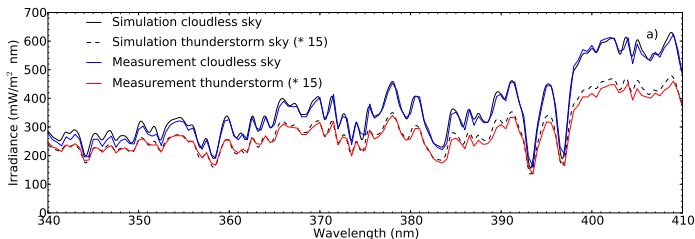


CIE UV dose for action spectrum  $A$

$$E_{\text{CIE}}^\downarrow(t) = \int_0^\infty A_{\text{CIE}}(\lambda) E^\downarrow(\lambda, t) d\lambda$$



# Measured quantities, UV spectrum



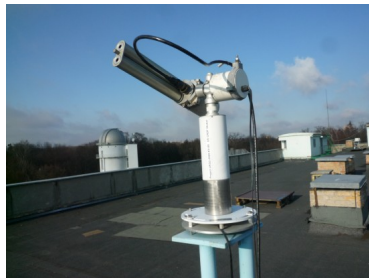
Surface based spectrometer measures downward irradiance as a function of wavelength.

$$E^{\downarrow}(\lambda, t) = \int_0^{2\pi} d\phi \int_0^1 \mu L(\lambda, \mu, \phi) d\mu$$

# Measured quantities, AERONET

Surface-based radiance band measurements to deduce aerosol information.

$$L_i(\theta, \phi, t) = \int_{\lambda_1}^{\lambda_2} R(\lambda)L(\theta, \phi, \lambda, t)d\lambda$$



# Measured quantities, SEVIRI

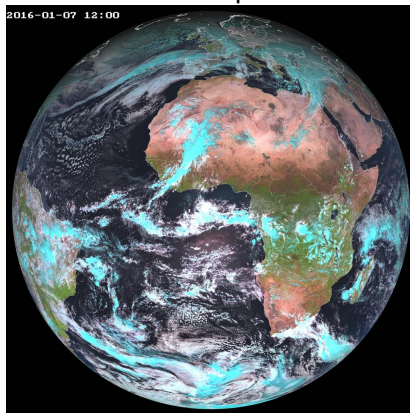
Spinning Enhanced Visible  
Infra-Red Imager (SEVIRI)

Measures radiances in band  $i$

$$L_i(\theta, \phi, t) = \int_{\lambda_1}^{\lambda_2} R(\lambda) L(\theta, \phi, \lambda, t) d\lambda$$

Instruments like MODIS, VIIRS etc.  
measures similar quantities.

RGB composite

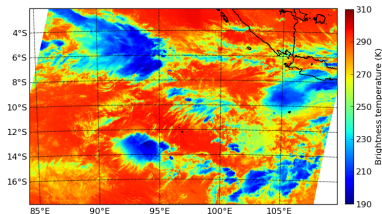




# Measured quantities, IASI and AVHRR

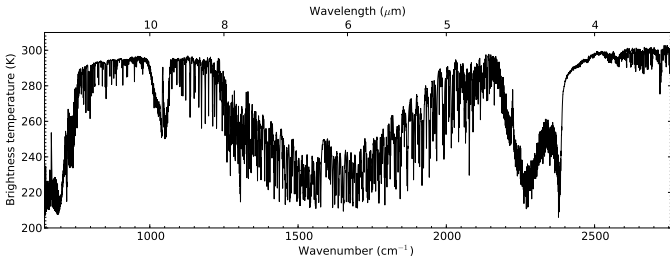
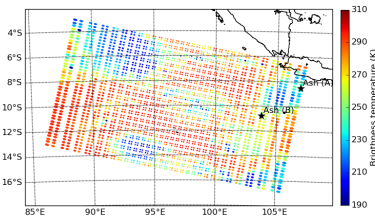
AVHRR

$$L_i(\theta, \phi, t) = \int_{\lambda_1}^{\lambda_2} R(\lambda) L(\theta, \phi, \lambda, t) d\lambda$$



IASI:

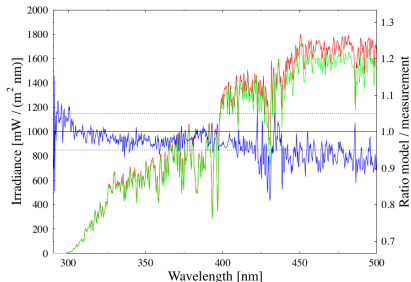
$$L(\theta, \phi, \lambda, t)$$



- Comparison with analytical solutions. DISORT has test code for this.
- Compare with other models
  - Need to agree on cross sections, atmospheric profile and layering.
  - What if all are wrong?
- Compare with measurements
  - Models do not compute what is being measured. Simulate instrument response.
  - Measurements do not precisely measure the wanted quantity. “Correct” measurements.
  - Models include approximations.
  - Measurements have uncertainties.
  - In atmospheric science measurements are generally not repeatable.
  - Closure experiments extremely difficult to make.

# Validation: Surface irradiance

Numerous studies in the UV and visible of model-measurement comparisons of surface and airborne UV and visible irradiances and actinic fluxes (Kylling et al., 1993; Mayer et al., 1997; Kylling et al., 1998; Van Weele et al., 2000; Bais et al., 2003; Hofzumahaus et al., 2002; Kylling et al., 2003; Wendisch and Mayer, 2003).

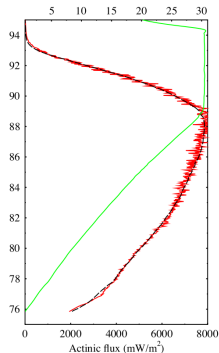
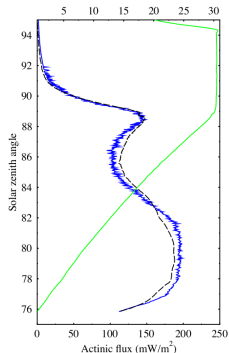
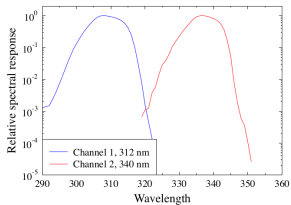


red line: measurements

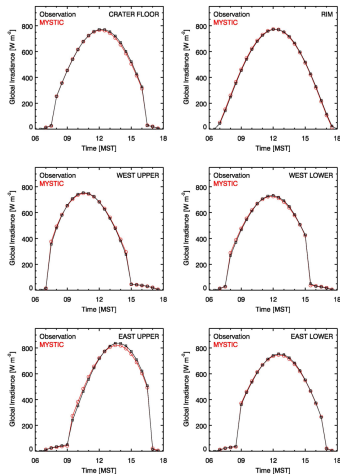
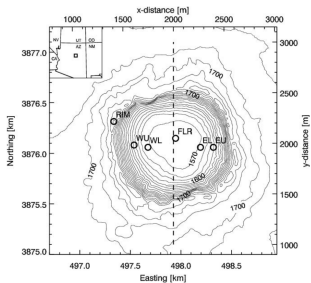
green line: model

blue line: model/measurement ratio

# Validation: balloon measurements



# Validation: 3D



From Mayer et al. (2010).

For 3D model intercomparison: [https://i3rc.gsfc.nasa.gov/I3RC-intro\\_new.php](https://i3rc.gsfc.nasa.gov/I3RC-intro_new.php)

# Today's exercises:

- Calculate solar spectrum, include filter function, integrate
- Calculate solar spectrum, include slit function
- Calculate thermal spectrum, include filter function, integrate
- Plot results

## Hints:

- **example input files:** `UVSPEC_LOWTRAN_SOLAR.INP`
- **options** `filter_function_file`, `slit_function_file`

# References I

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